

# Remote sensing of the planets

A vast literature and methodology exists in optical and radar astronomy that parallels and often exceeds our methods used in remote sensing



Space –  
the final  
RS  
frontier

# Welcome to the Planets<sup>™</sup>



Mercury



Venus



Earth



Mars



Jupiter



Saturn



Uranus



Neptune



Pluto



Small Bodies

Planet  
Profiles

Explorers

Glossary

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What's  
New

# Methods and wavelengths used on planetary missions

METHOD	EM SPECTRUM	INFORMATION	INTERPRETATION	MISSION
Gamma-Ray Spectroscopy	Gamma rays	Gamma spectrum	K, U, Th Abundances	Apollo 15, 16; Venera
X-ray Fluorescence spectrometry	X-rays	Characteristic Wavelengths	Surface mineral/ chemical comp.	Apollo; Viking Landers
Ultraviolet Spectrometry	UV	Spectrum of Reflected sunlight	Atmospheric Composition: H, He, CO <sub>2</sub>	Mariner; Pioneer; voyager
Photometry	UV, Visible	Albedo	Nature of Surface; Composition	Earth Telescopes; Pioneer
Multispectral Imagers	UV, Visible, IR	Spectral and Spatial	Surface Features; Composition	On most missions
Reflectance Spectrometers	Visible, IR	Spectral intensities of reflected solar radiation	Surface Chemistry; mineralogy; processes	Telescopes; Apollo
Laser Altimeter	Visible	Time delay between emitted and reflected pulses	Surface Relief	Apollo 15, 16, 17
Polarimeter	Visible	Surface Polarization	Surface Texture; Composition	Pioneer; Voyager

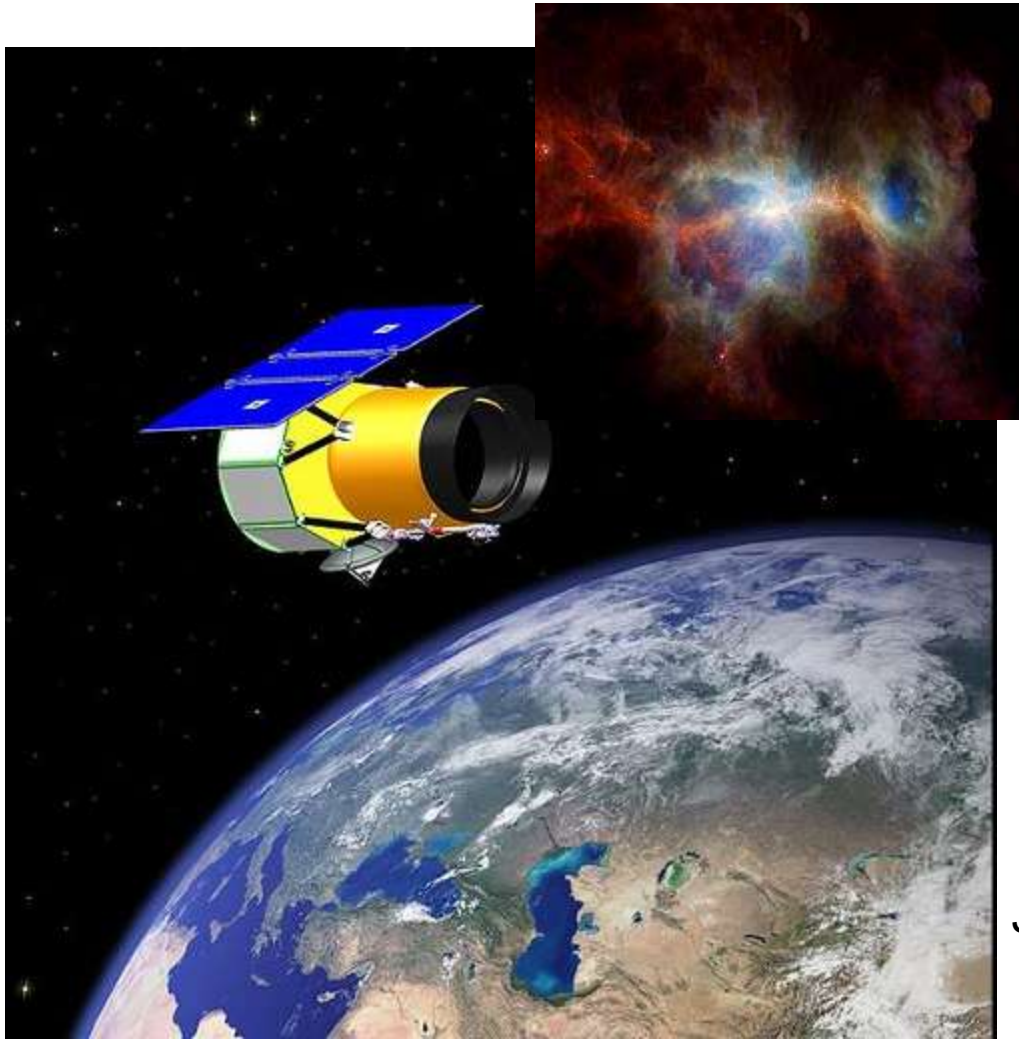
Infrared Radiometer (includes scanners)	Infrared	Thermal radiant intensities	Surface and atmospheric temperatures; compos.	Apollo; Mariner; Viking; Voyager
Microwave Radiometer	Microwave	Passive microwave emission	Atmosphere/Surface temperatures; structure	Mariner; Pioneer Venus
Bistatic Radar	Microwave	Surface reflection profiles	Surface Heights; roughness	Apollo 14,15,16; Viking
Imaging Radar	Microwave	Reflections from swath	Topography and roughness	Magellan; Earth systems
Lunar Sounder	Radar	Multifrequency Doppler Shifts	Surface Profiling and imaging; conductivity	Apollo 17
S-Band Transponder	Radio	Doppler shift single frequency	Gravity data	Apollo
Radio Occultation	Radio	Frequency and intensity change	Atmospheric density and pressure	Flybys and Orbiters

\* Adapted from Billy P. Glass, *Introduction to Planetary Geology*, 1982, Cambridge University, Press

Hyperspectral    visible-NIR    Reflected solar radiation    Surface    Mars



**Wide-field Infrared Survey Explorer (WISE) since Nov 20, 2009 ..**  
**Succeeded by NEOWISE (2013) Looking out into space**  
Detectors at 3.4, 4.6, 12 and 22 microns, chilled to 10 Kelvin



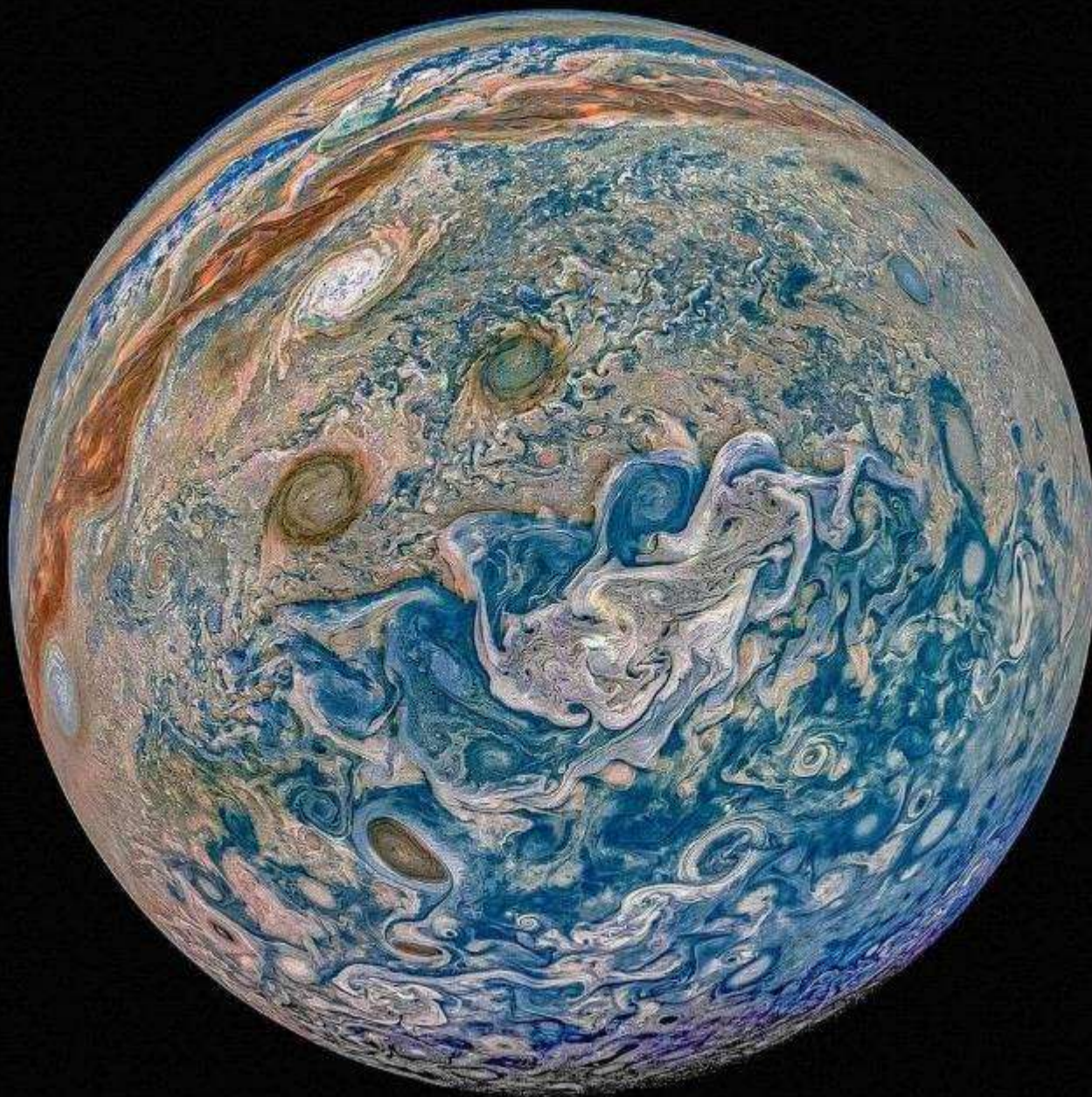
**Hubble telescope 1990**  
0.1 – 0.8 microns, 535km



**James Webb Space Telescope**  
0.6 – 28 microns, 1.5 million km  
Launched Dec 25, 2021

[http://www.nasa.gov/mission\\_pages/WISE/main/index.html](http://www.nasa.gov/mission_pages/WISE/main/index.html)





Images of Jupiter

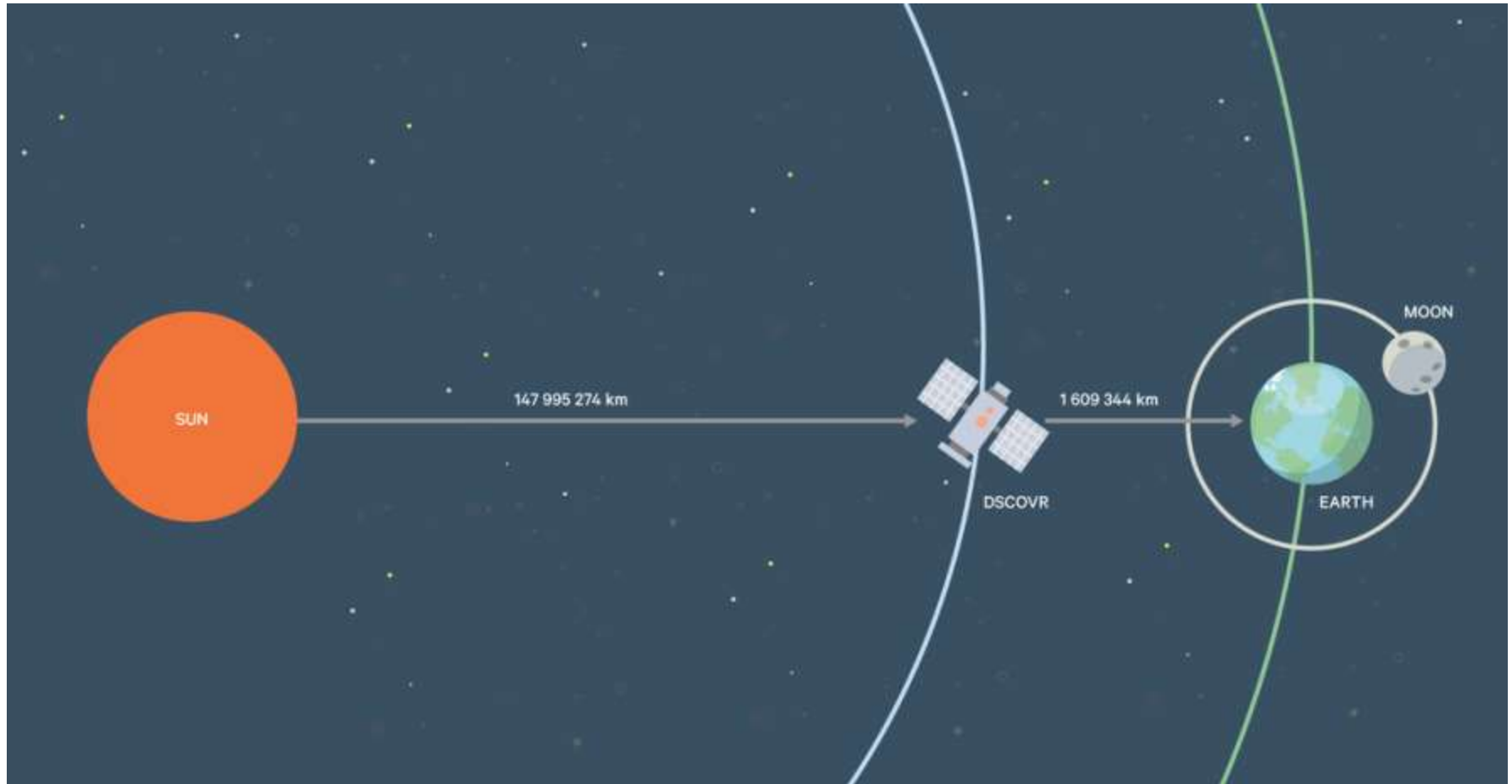
JWST

Hubble



# **DSCOVER: Deep Space Climate Observatory - 2015**

**1 million miles away – ~10km resolution**  
**– orbiting at ‘Lagrange point L1 = gravitational pull**



Proposed by Al Gore, 1998 to study earth and solar wind  
The first satellite orbiting in deep space ..... ‘Goresat’

# 1. EPIC: Earth Polychromatic Imaging Camera, 10 bands

EPIC Wavelengths and main data products

Wavelength(nm)	Full Width (nm)	Primary Application
317.5 ± 0.1	1 ± 0.2	Ozone, SO <sub>2</sub>
325 ± 0.1	2 ± 0.2	Ozone
340 ± 0.3	3 ± 0.6	Ozone, Aerosols
388 ± 0.3	3 ± 0.6	Aerosols, Clouds
443 ± 1	3 ± 0.6	Aerosols, Clouds
551 ± 1	3 ± 0.6	Aerosols
680 ± 0.2	3 ± 0.6	Aerosols, Vegetation
687.75 ± 0.2	0.8 ± 0.2	Aerosols, Vegetation, Clouds
764 ± 0.2	1 ± 0.2	Cloud Height
779.5 ± 0.3	2 ± 0.4	Clouds, Vegetation

2. Radiometer to measure radiance UV-TIR, - monitor earth temperature



# **‘Dark side’ of the Moon crossing Earth from DSCOVR satellite**

*daily images from EPIC*

Earth Polychromatic Imaging Camera (EPIC)

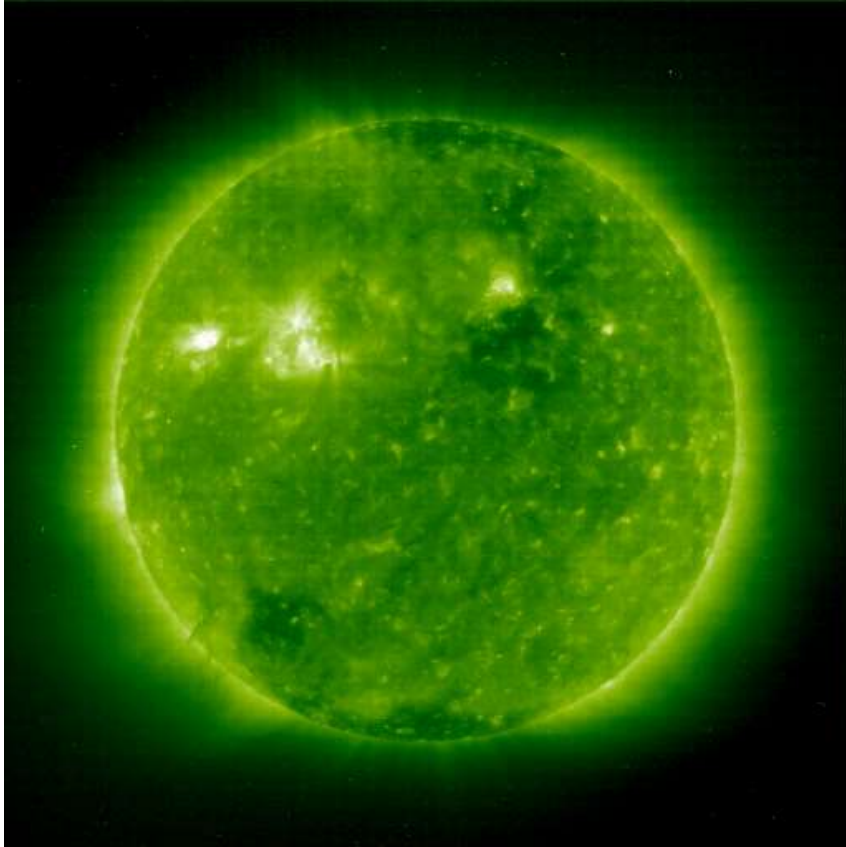
<http://epic.gsfc.nasa.gov/>



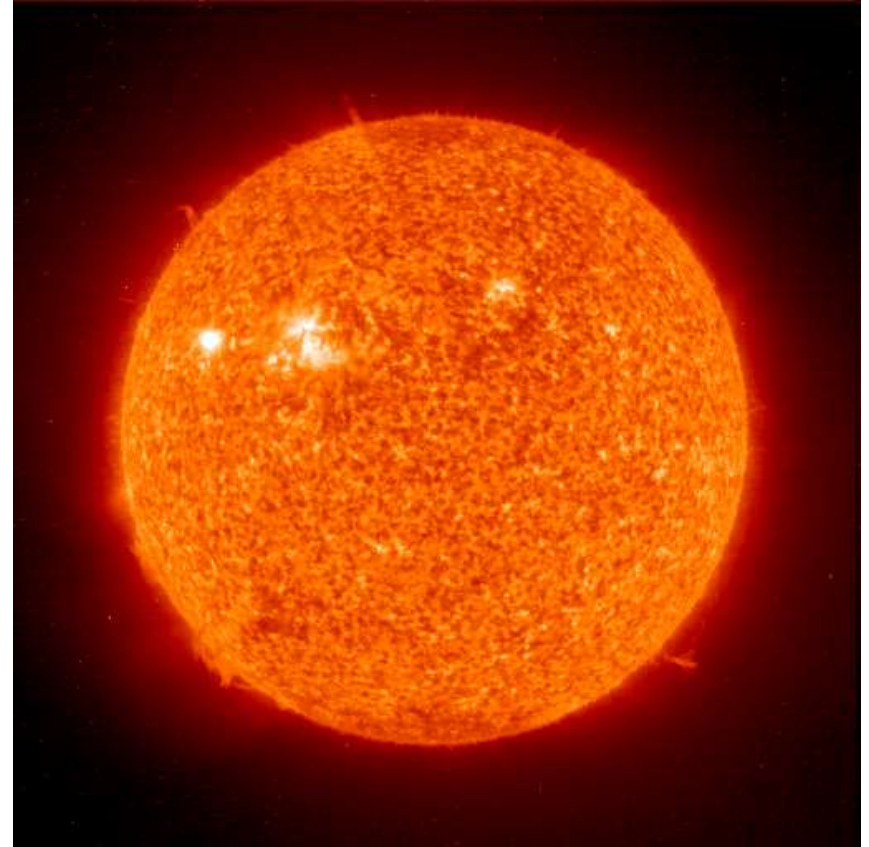
# SOHO the Solar & Heliospheric Observatory

.. is a project of international collaboration between ESA and NASA to study the Sun from its core to the outer corona and the solar wind.

Nov 18, 2009: 195 nm



304 nm (also 171 and 284)



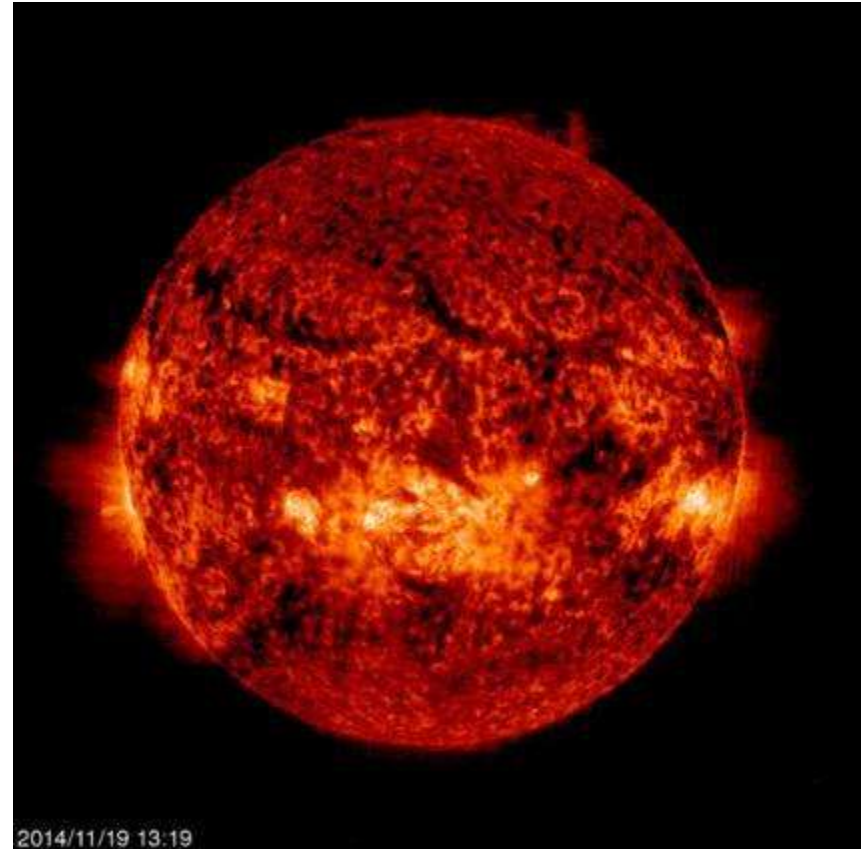
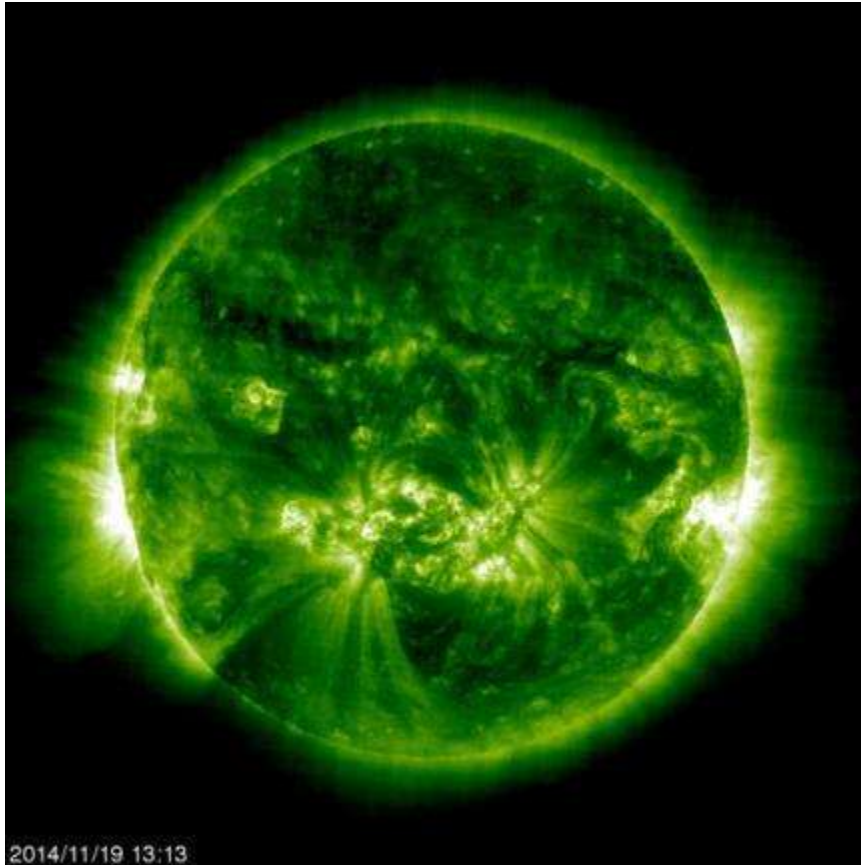
1995, cost €1 billion; Sensor: Extreme ultraviolet Imaging Telescope (EIT)

Located at sun-earth L1 gravitational orbit

# SOHO, the Solar & Heliospheric Observatory

.. is a project of international collaboration between ESA and NASA to study the Sun from its core to the outer corona and the solar wind.

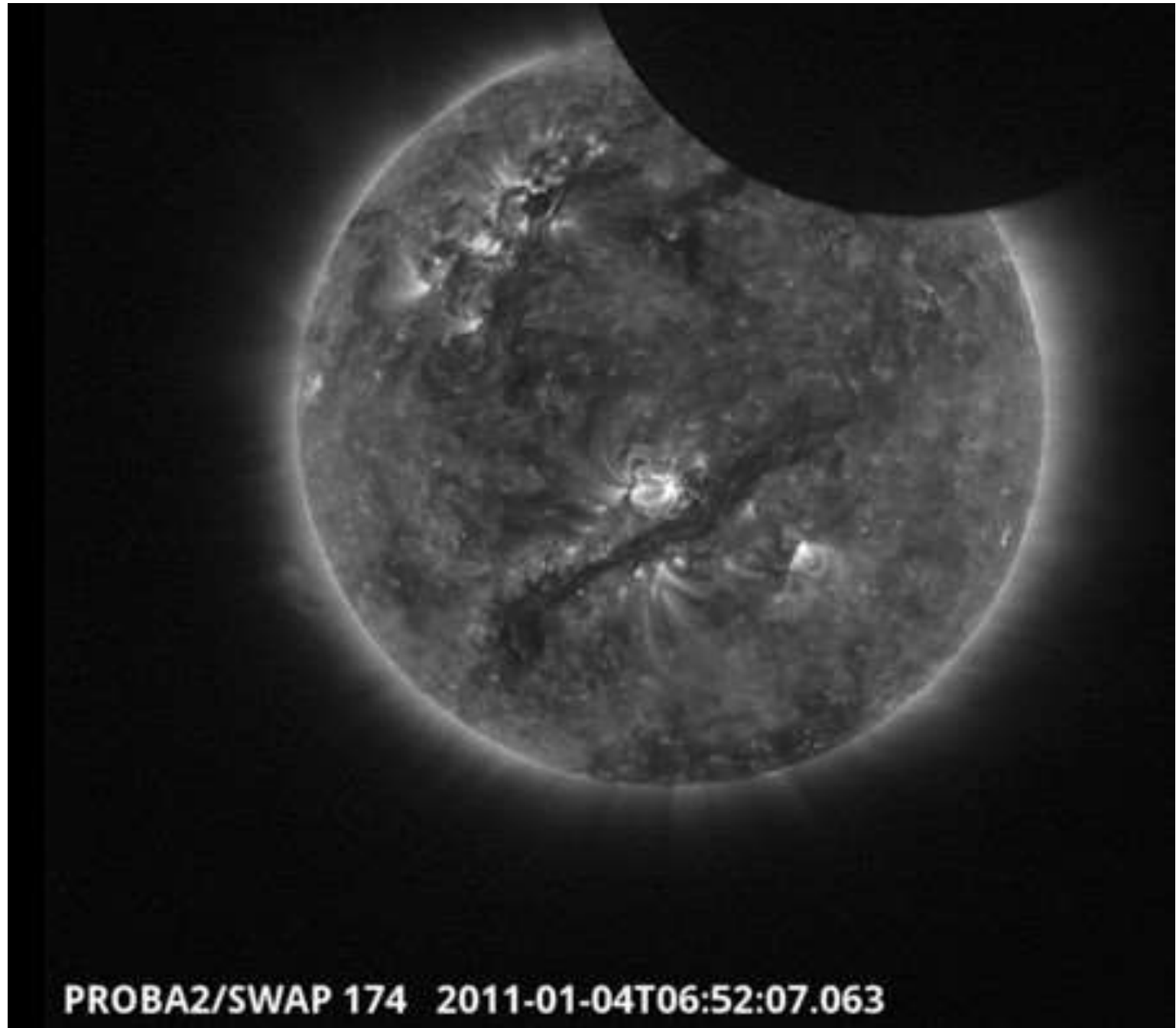
Later images - Nov 19, 2014: 195 nm      304 nm (also 171 and 284)



Daily images: <https://umbra.nascom.nasa.gov/newsite/images.html>



# Partial solar eclipse by the Moon



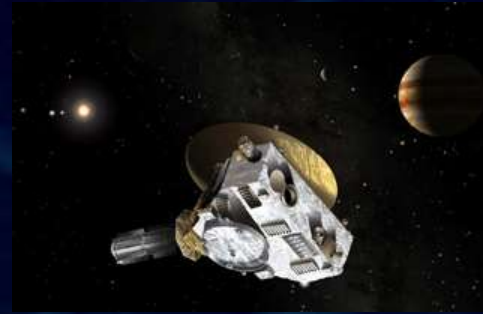
PROBA2/SWAP 174 2011-01-04T06:52:07.063

# Types of Remote Sensing Missions

- Fly-bys



Mariner Missions,  
Mercury and Venus



New Horizons 2007,  
Jupiter and Pluto

- Orbiters

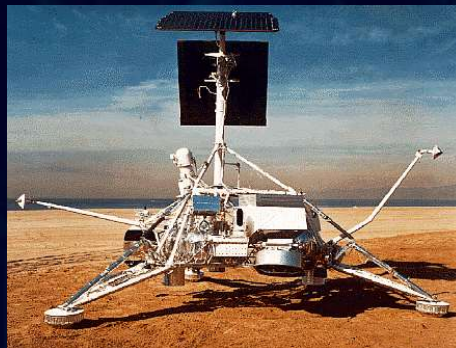


LandSat, Earth



Mars Reconnaissance Orbiter

- Landers /  
Rovers

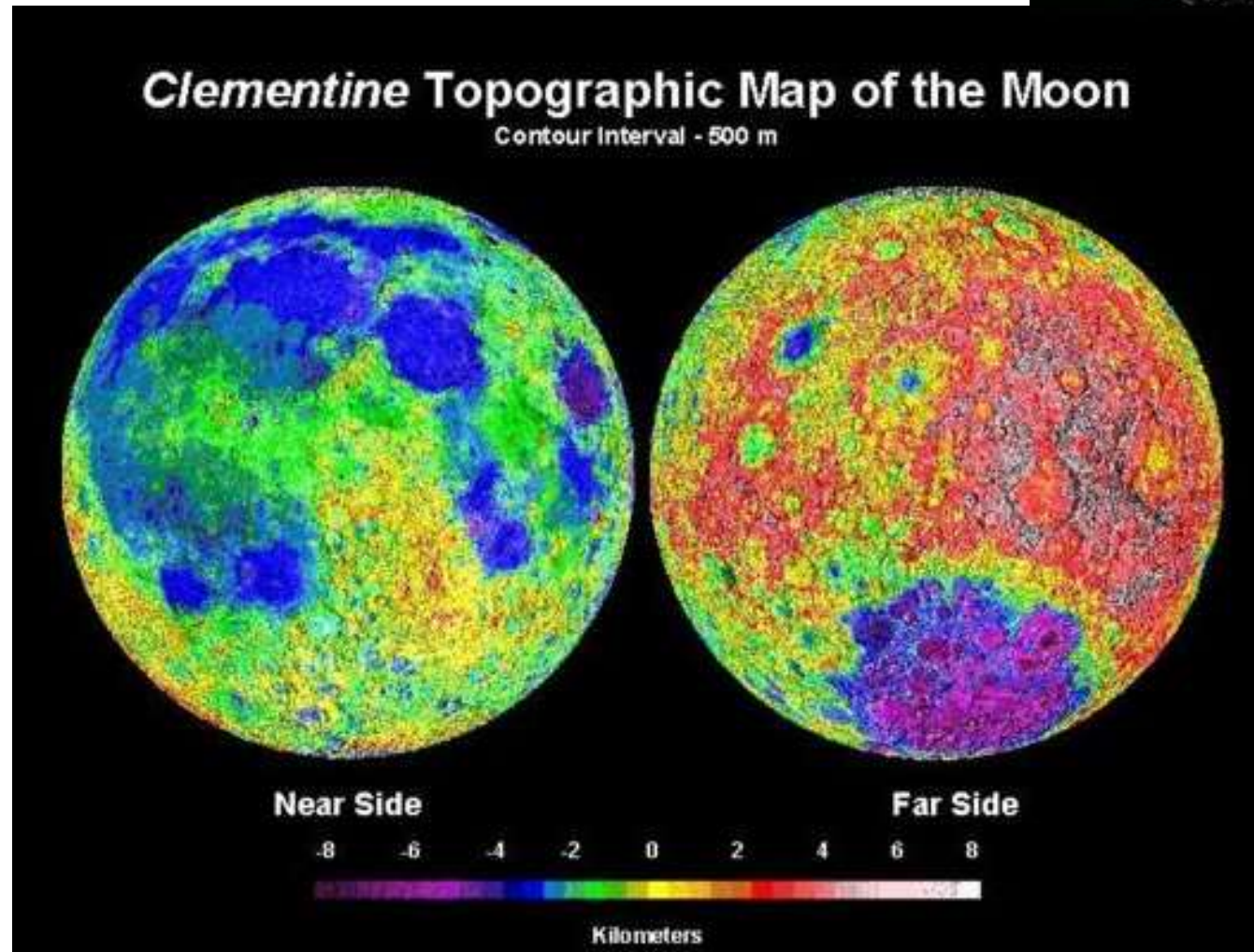


Surveyor Lunar Lander, 1966



Mars Rovers, 2003

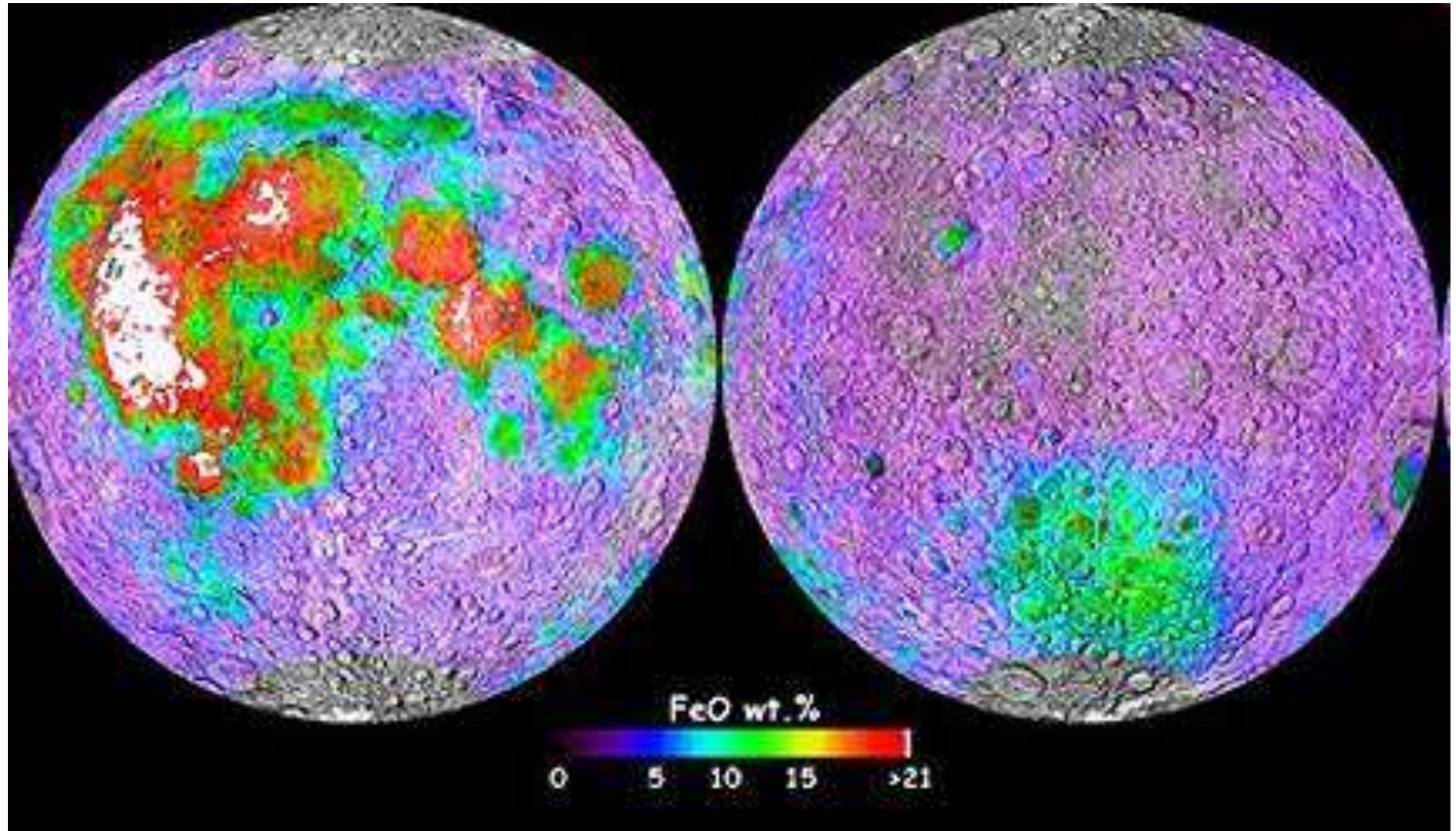
After Apollo (1972), the Moon was not revisited until an unmanned spacecraft, Clementine orbited to conduct mapping studies February 19 - April 21, 1994, using UV/Visible, Near IR, and Lidar



<https://www.google.ca/moon/>



specialized products include detailed maps of lunar topography and the distribution of several chemical elements, such as iron (Fe) and titanium (Ti), determined by analyzing reflectance variations at 0.75  $\mu$ m and 0.95  $\mu$ m, where these elements absorb radiation.





In mid-April 2000, the Terra spacecraft was turned upside down and pointed at the Moon. This ASTER image was acquired at that time, showing band 3 (NIR) in black and white.



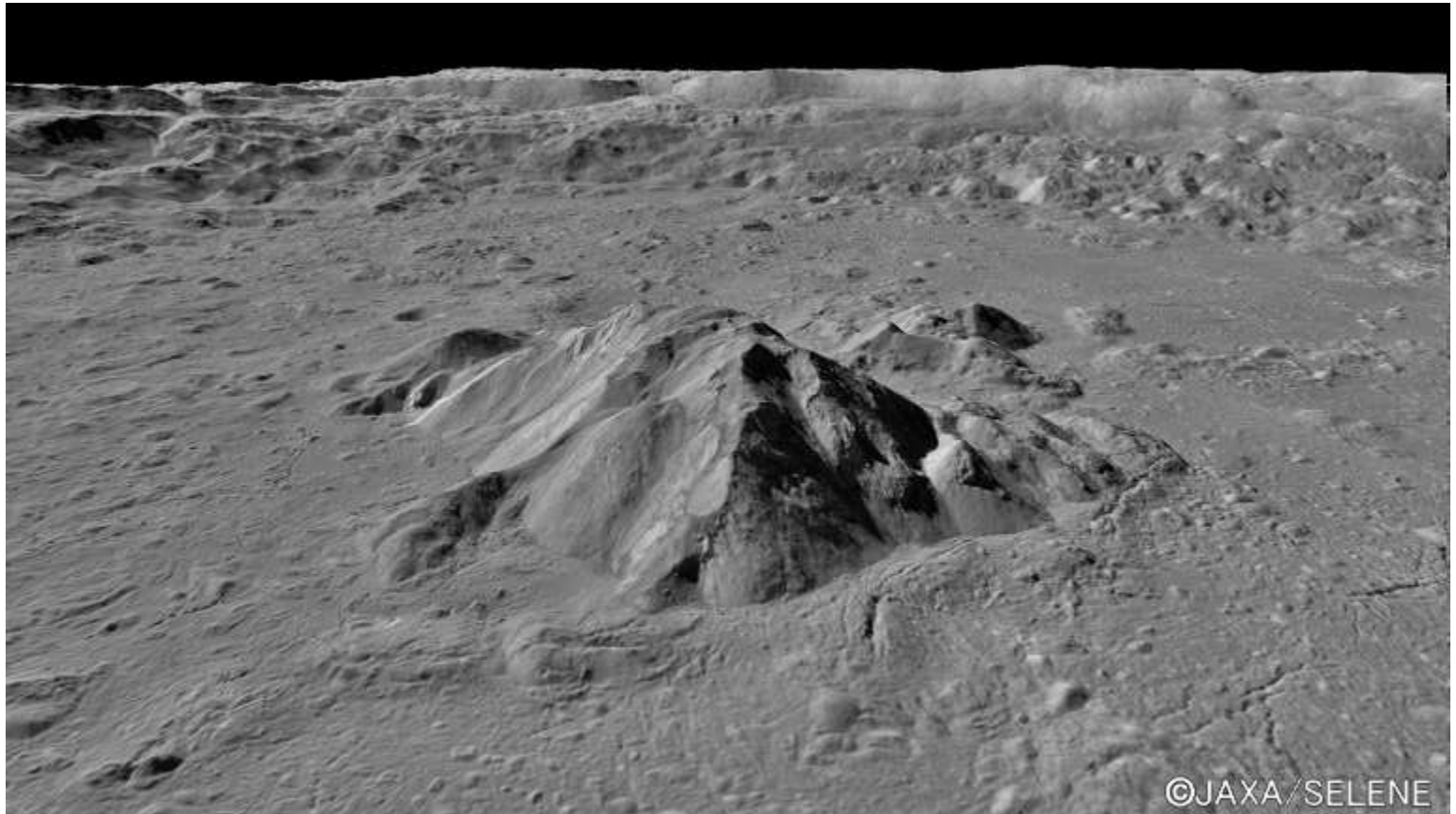
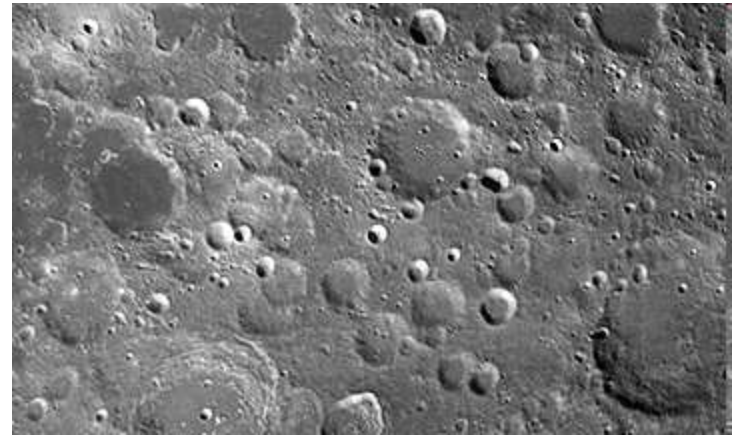
# Recent Lunar missions

Right: China: Chang'e-1 (2007)

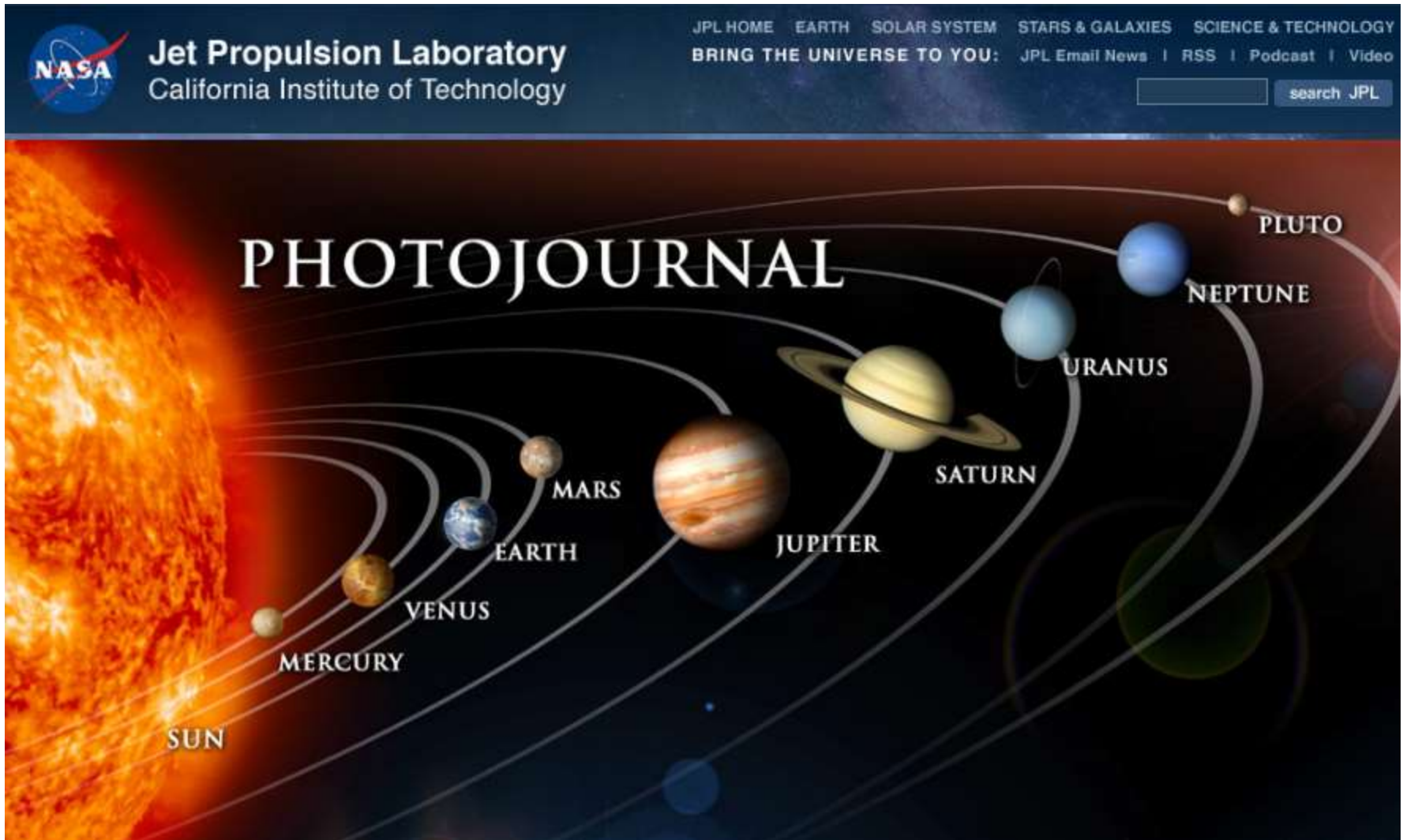
Chang'e-3 (2013) soft landed on moon

Below: Japan - Kayuga

Launched 2007, impacted on lunar surface 2009 (near South Pole)





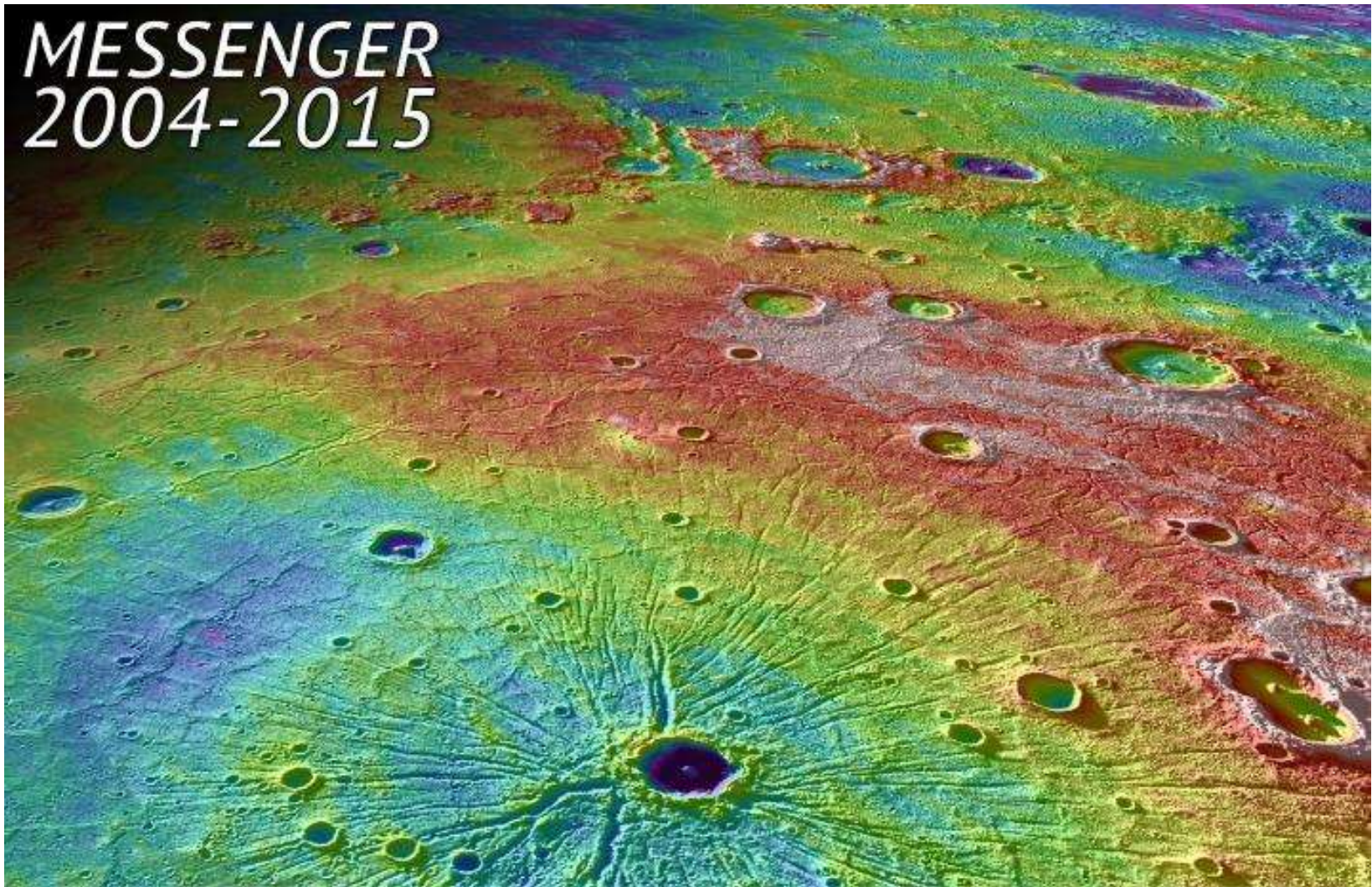


satellite links: Sun 186 Mercury 1315, Venus 153, Earth 2522, Mars 12216,  
Jupiter 1151, Saturn 3373, Uranus 61, Neptune 89, Dwarf Planets e.g. Pluto 817



# **Mercury Messenger:**

**Mercury Surface, Space Environment, Geochemistry, & Ranging**



**Mercury Dual Imaging System (MDIS) and Mercury Laser Altimeter (MLA)**



# Venus

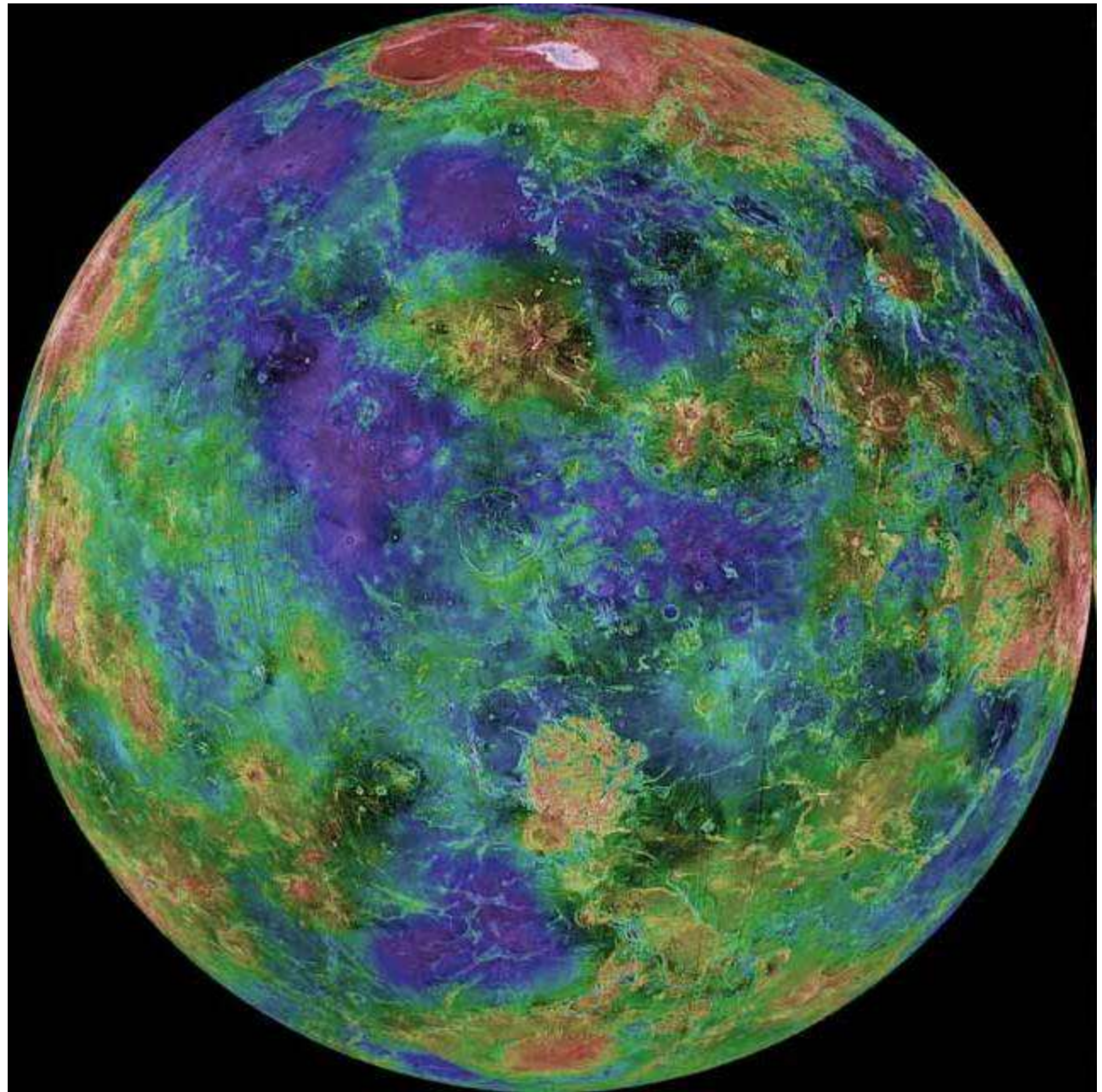
From Magellan

-first imaging device  
launched from  
Shuttle 1989

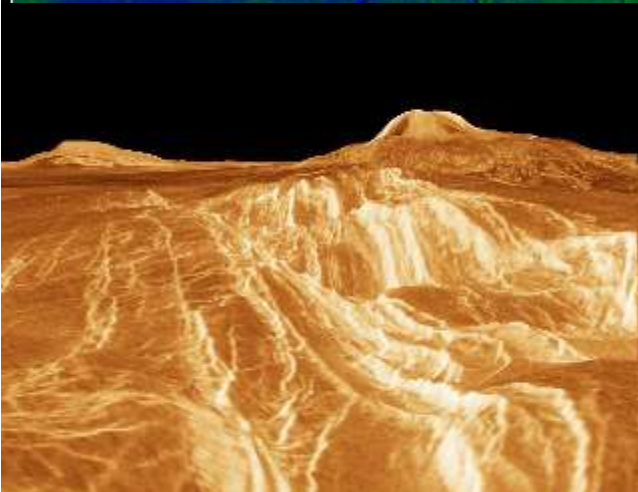
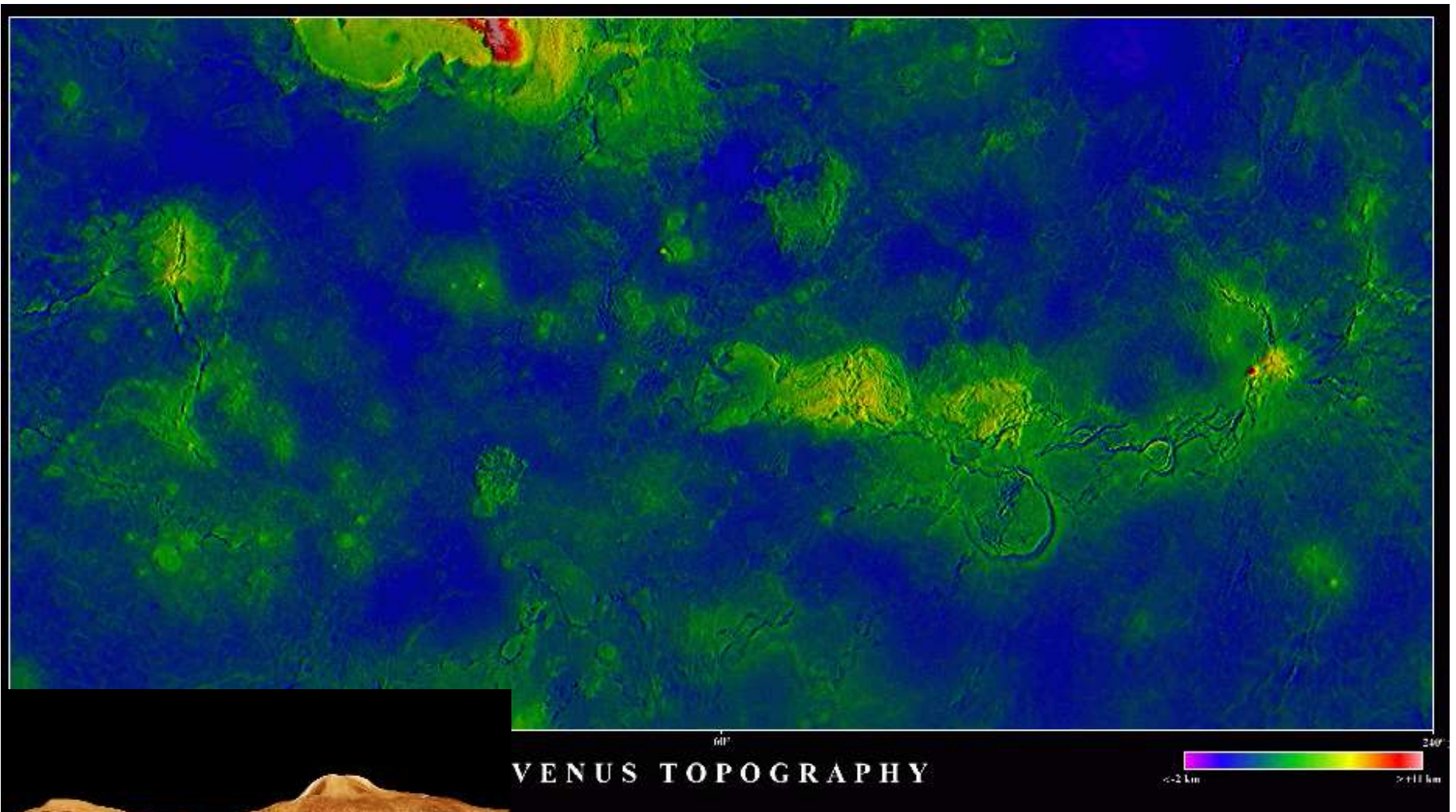
Planet is Cloud  
covered

Radar 100m

Composite colours  
based on elevations







<http://www.solarviews.com/eng/venus.htm>

# Mars Global Surveyor (1996) Instruments



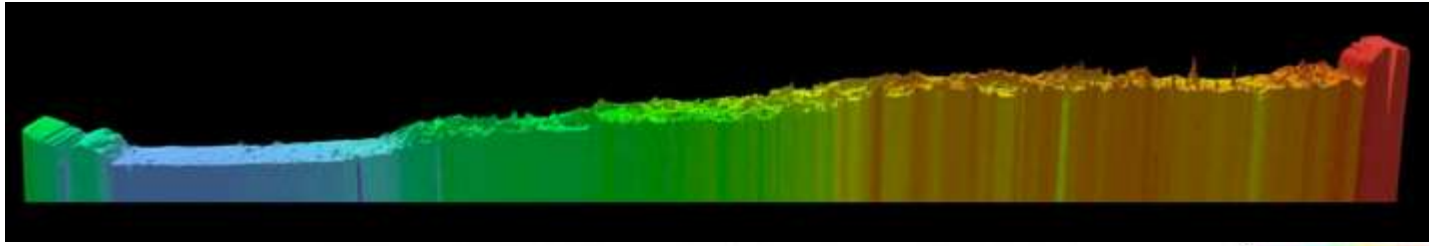
MOLA - Mars Orbiter Laser Altimeter

MOC - Mars Orbiter Camera

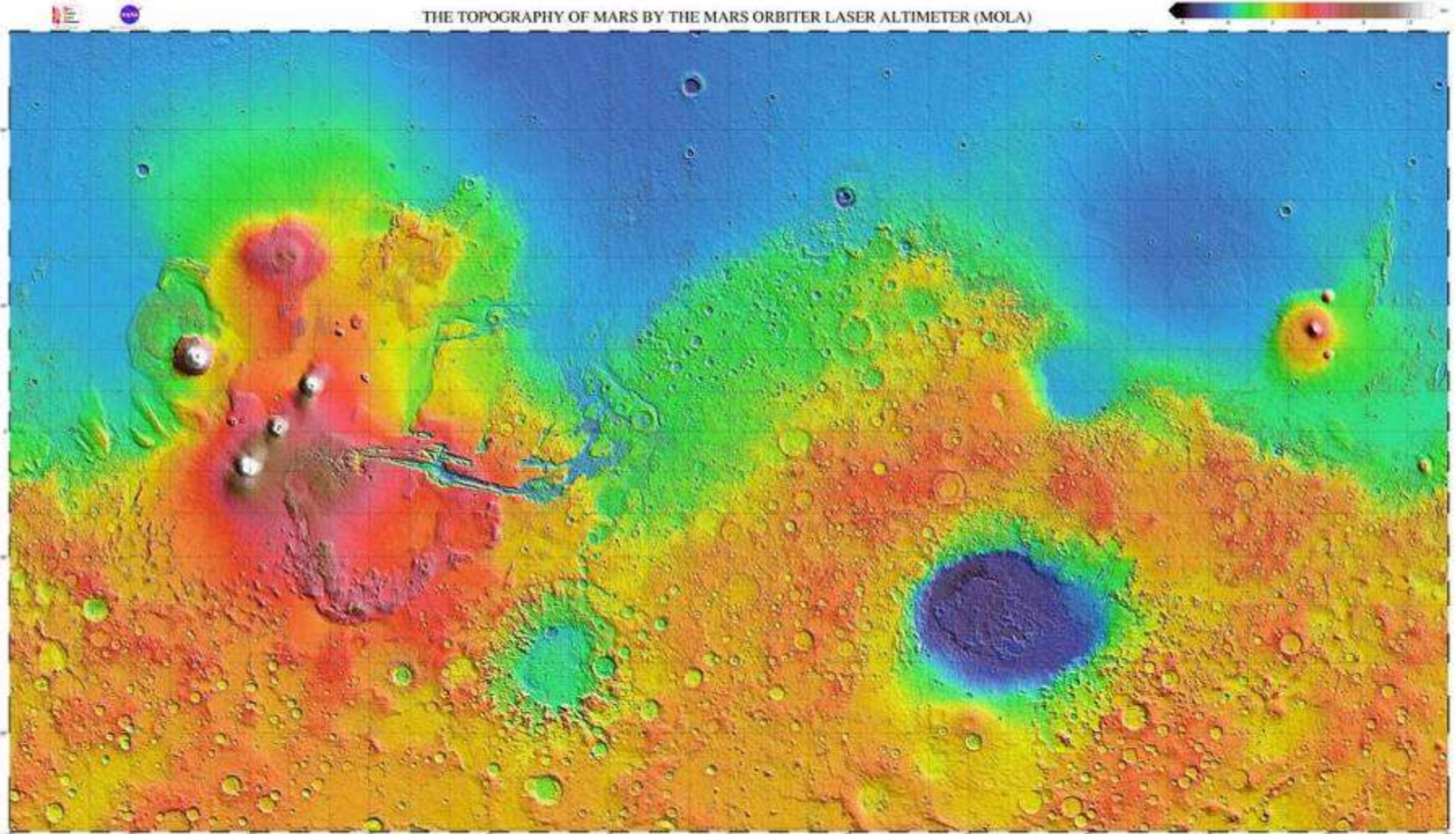
TES - Thermal Emissions



DEM resolution in z = 30cm! (N. Pole to S. Pole transect)



THE TOPOGRAPHY OF MARS BY THE MARS ORBITER LASER ALTIMETER (MOLA)



<https://www.google.ca/mars/>

Elevation, Panchromatic, Thermal





**Suspected rock glacier, Mars Orbiter Camera JPL/NASA**

**Resolution = 1m**

- **MOC has produced over 250,000 images to 2020**

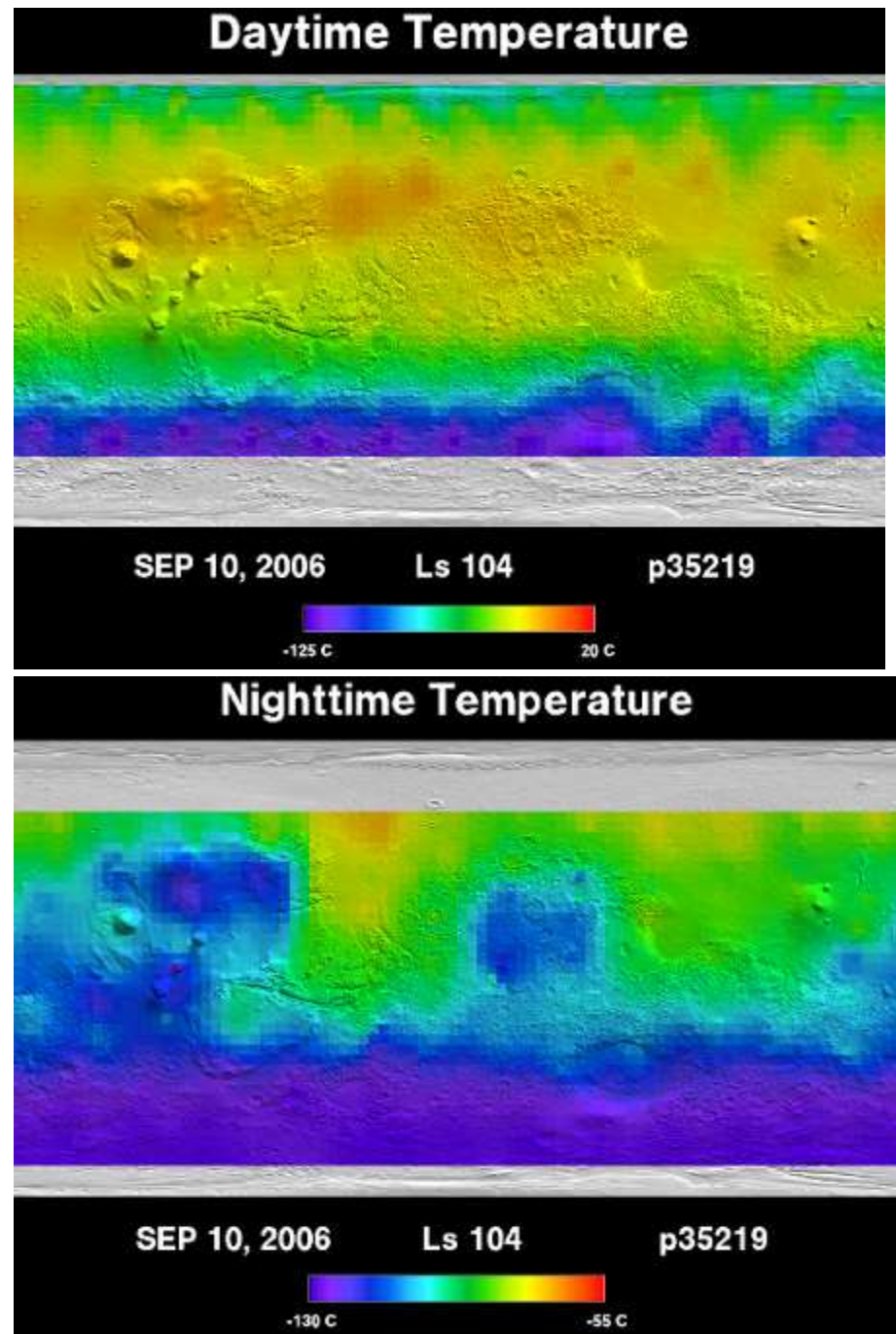
# Thermal Emission Spectrometer

6 to 50 ( $\mu\text{m}$ ),

143 bands

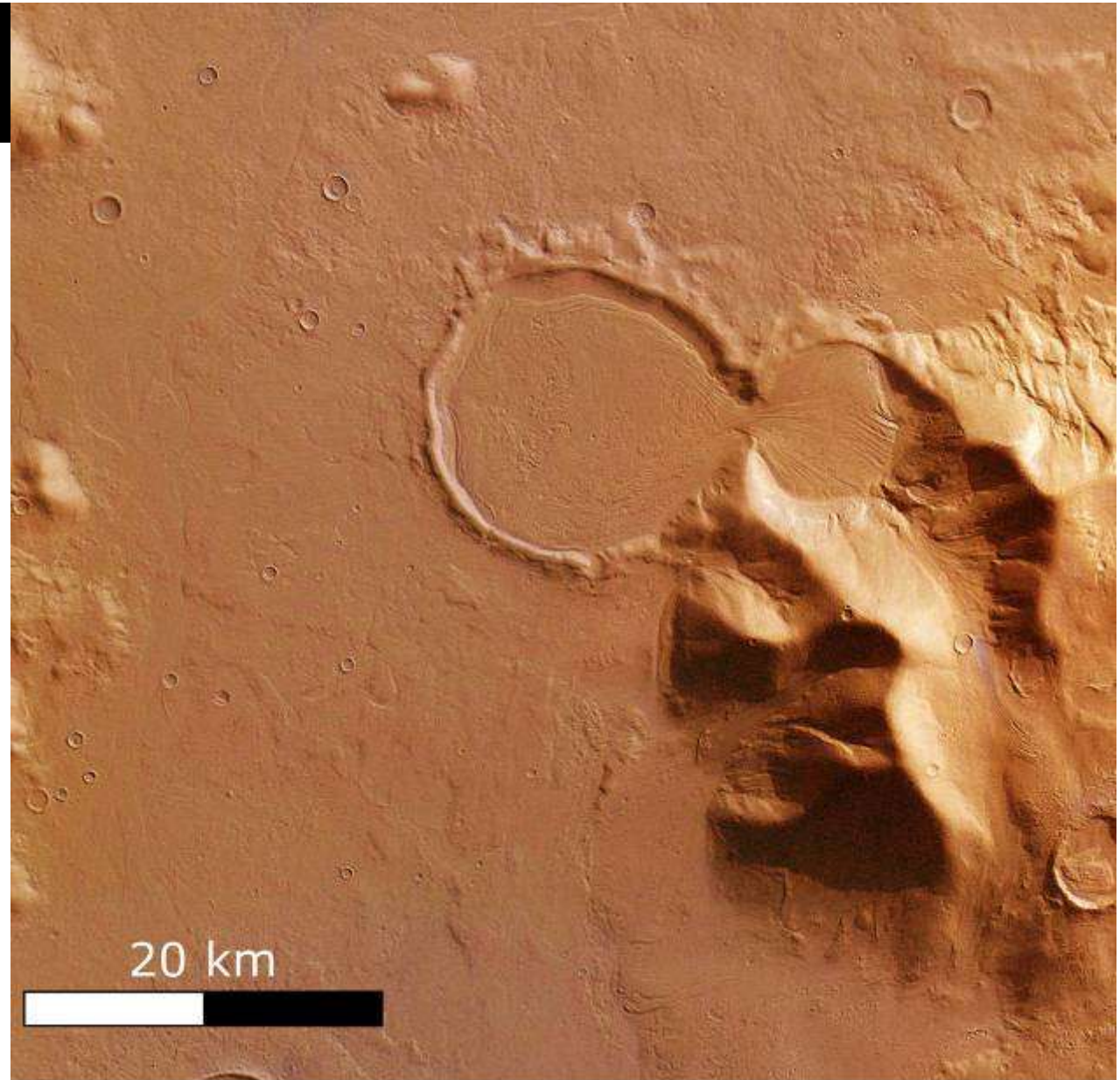
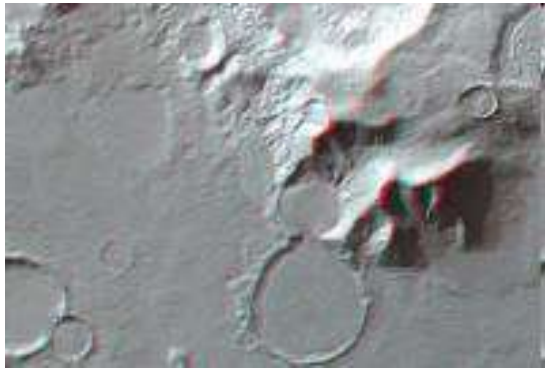
Onboard Mars Global  
Surveyor

1996–2006



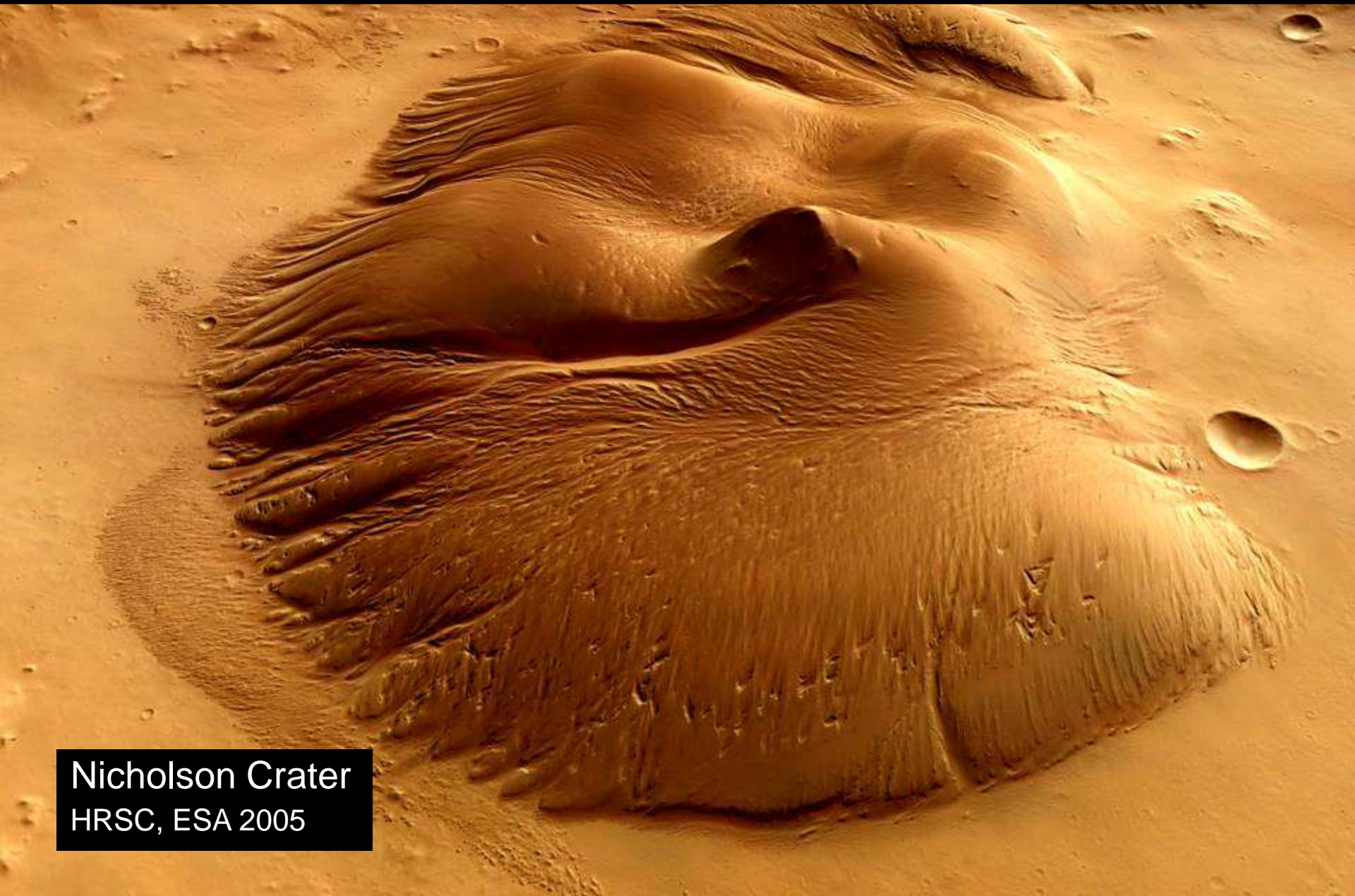
# Mars Express (ESA, 2003): High Resolution Stereo Camera Resolution 2-10m

The “hourglass” feature  
HRSC, ESA





# Mars Express: High Resolution Stereo Camera



Nicholson Crater  
HRSC, ESA 2005

# Mars Reconnaissance Orbiter (2005)

## Onboard:

- **HiRISE** - High Resolution Imaging Science Experiment (Visible and infrared wavelengths)
- **CRISM** - Compact Reconnaissance Imaging Spectrometer for Mars
- **CTX** - Context Imager  
Takes low resolution overview images for geological context





# Mars Reconnaissance Orbiter: HiRISE

2005-

**MRO HIGH RESOLUTION IMAGING SCIENCE EXPERIMENT (HIRISE) -1 foot (0.3m)**  
three bands, 400–600 nm ([blue-green](#)), 550–850 nm ([red](#)) 800–1,000 nm ([near infrared](#))



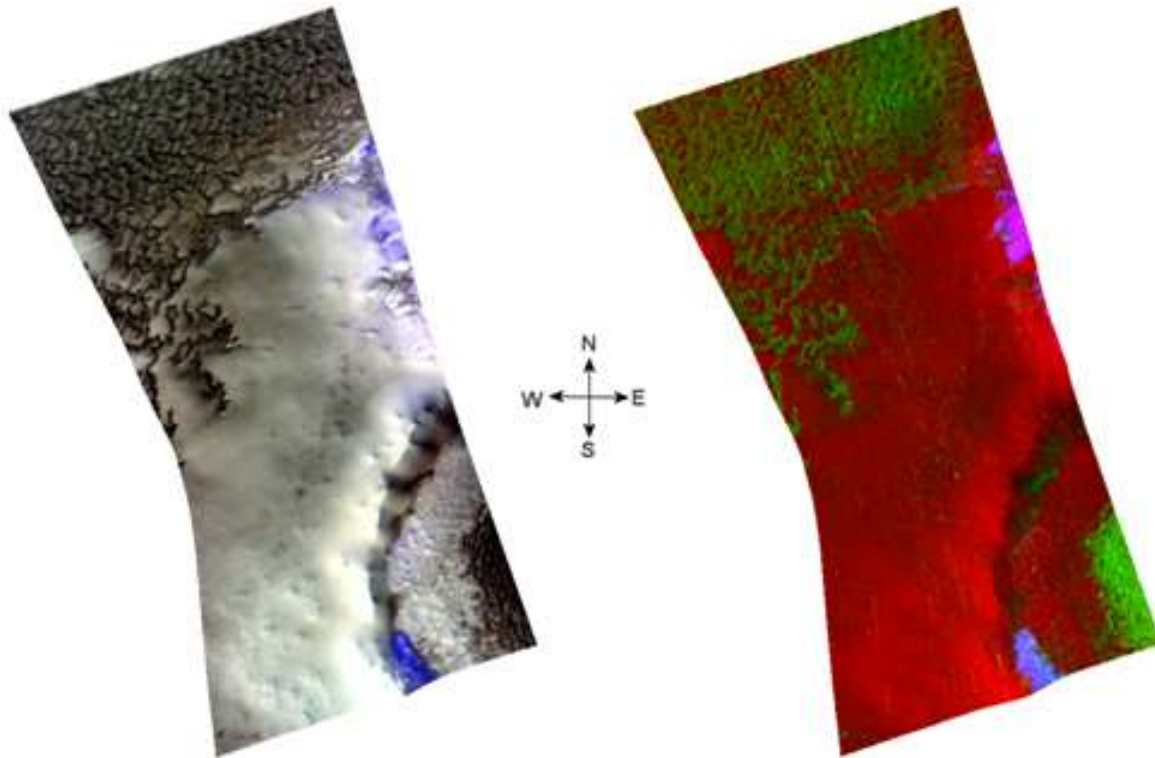
<http://hirise.lpl.arizona.edu/nea.php>

[http://marsoweb.nas.nasa.gov/HiRISE/hirise\\_images/](http://marsoweb.nas.nasa.gov/HiRISE/hirise_images/)



# Compact Reconnaissance Imaging Spectrometer for Mars (CRISM) on NASA's Mars Reconnaissance Orbiter (2005)

CRISM covers wavelengths from 0.362 to 3.92 microns (362 to 3920 nm) at 6.55 nanometers/channel, to identify a broad range of minerals on the Martian surface.



**False colour infrared: Red = dust, blue = water ice**

**Green = polyhydrated sulphate,**

# Thermal Emission Imaging System (THEMIS) 2001

This is a special camera on the Mars Odyssey spacecraft (2001). Its main tasks are mapping rock mineralogies and detecting heat, which yields information on the Martian surface.

**THEMIS is a multi-wavelength camera**

**5 visual bands:**

0.425 microns, 0.540 microns, 0.654 microns,  
0.749 microns, 0.860 microns

**10 infrared bands:**

6.78 microns (used twice), 7.93 microns, 8.56  
microns, 9.35 microns, 10.21 microns, 11.04  
microns, 11.79 microns, 12.57 microns, 14.88  
microns

**Resolution:**

visual images, 59 feet (18 meters) per pixel

infrared images, 328 feet (100 meters) per pixel

<http://themis.asu.edu/gallery>



# MARS ODYSSEY MISSION

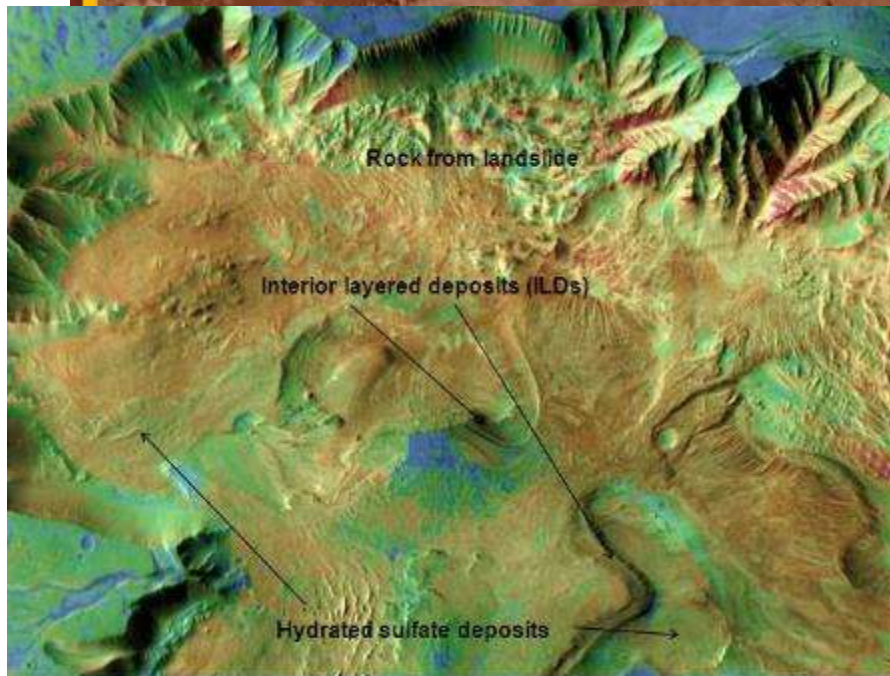
# THEMIS

THERMAL EMISSION IMAGING SYSTEM



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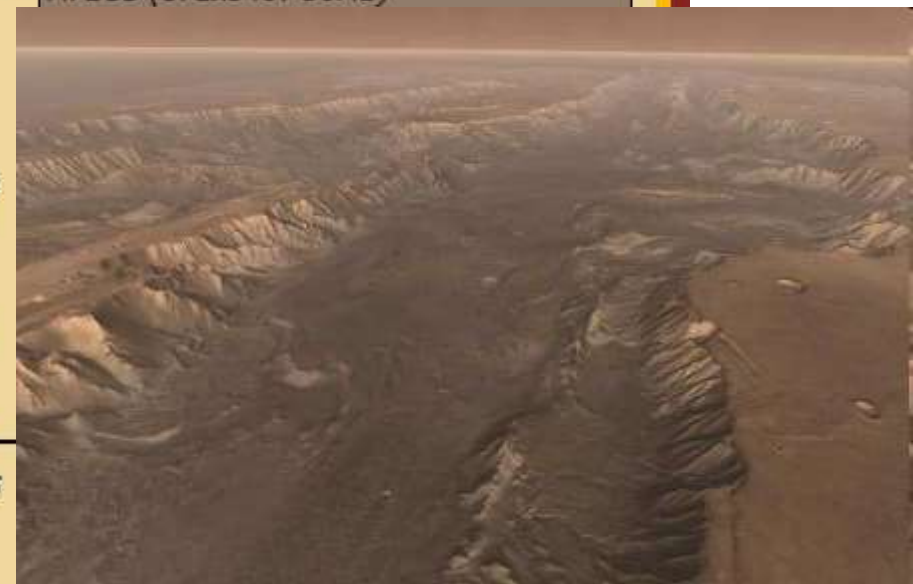
## *Flight Into Mariner Valley: The Movie*



### **Watch the Movie**

[Watch on Google Video \(low bandwidth\)](#)

MPEG1 (872x540: 56MB)



*The Grandest Canyon of all isn't on Earth, it's on the planet Mars - Valles Marineris, or Mariner Valley.*

[http://themis.asu.edu/valles\\_video](http://themis.asu.edu/valles_video)

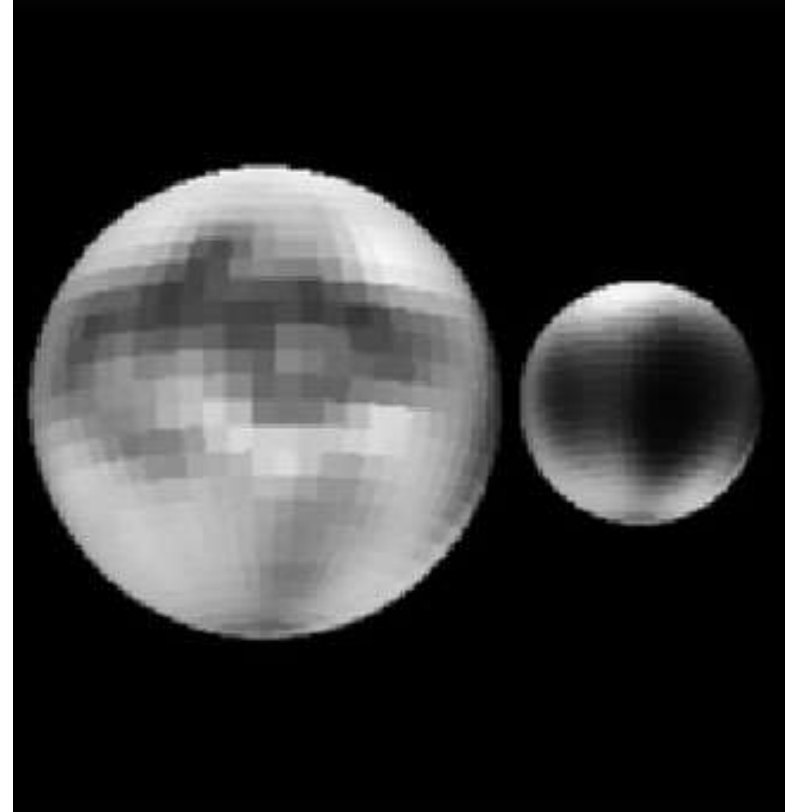
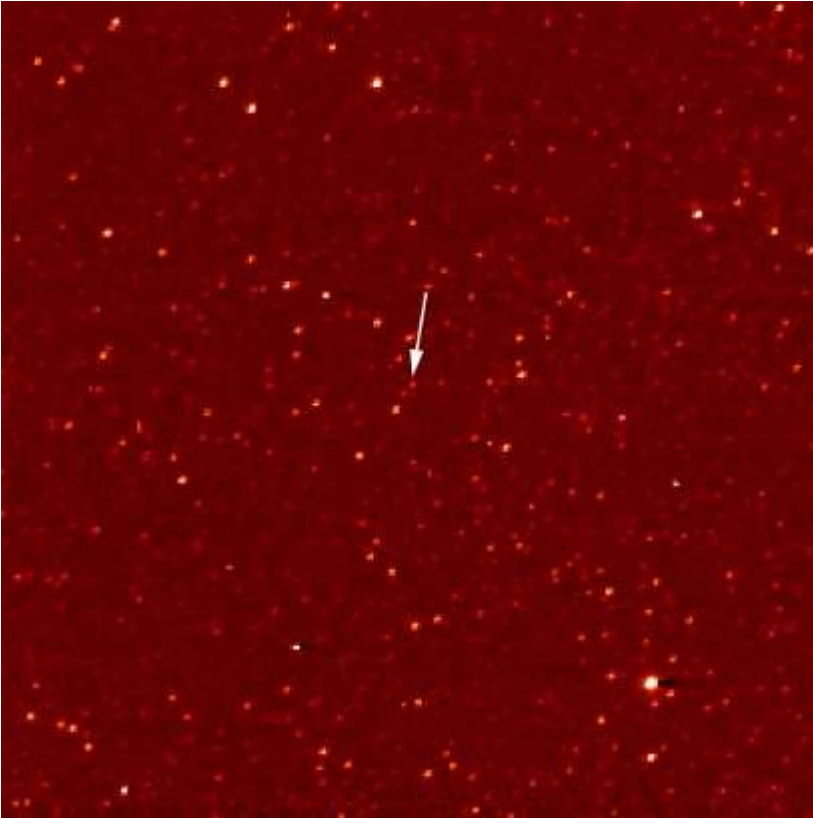
<http://themis.asu.edu/maps>

<http://mars.nasa.gov/mars3d>

Mars Exploration Rover Missions 2004



# Pluto and Charon (pre-2015)

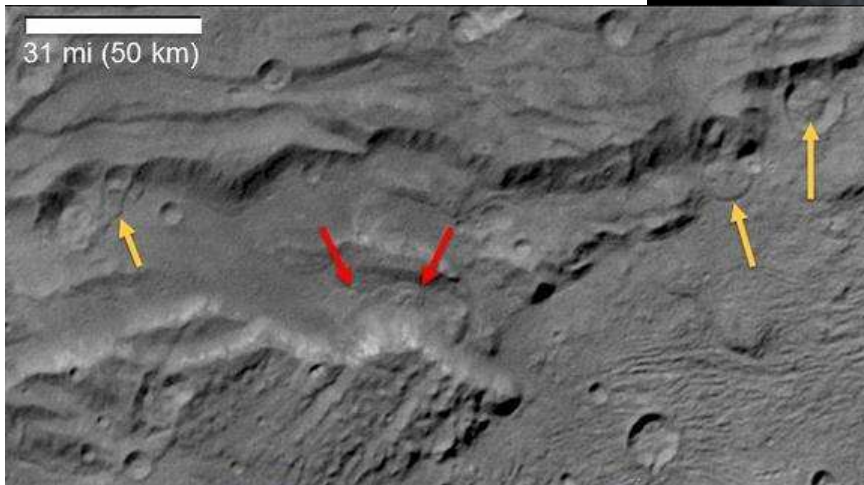


# New Horizons Mission 2015

Launched 2006

**Alice** Ultraviolet sensor and  
**Ralph** ('The Honeymooners')

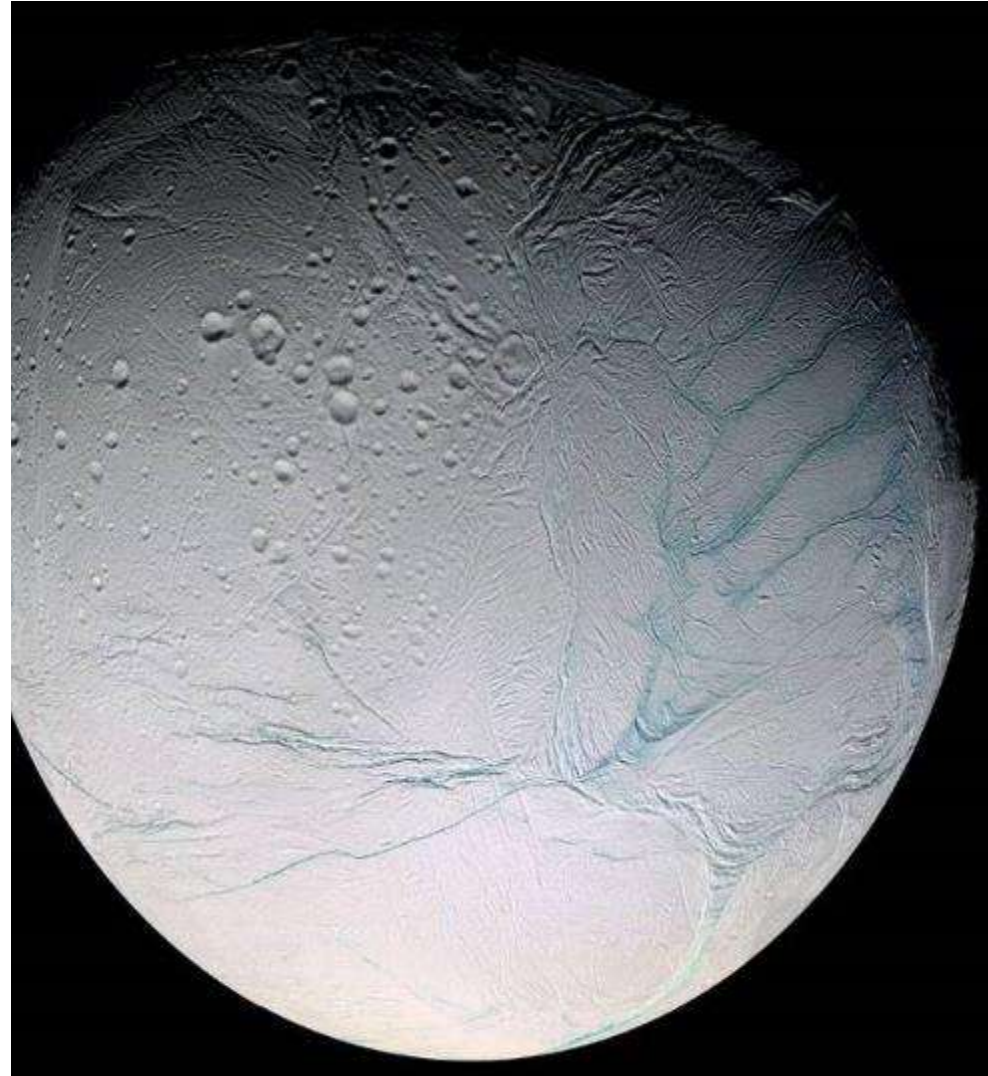
## Landslides on Charon



## The Planets and Their Moons

PLANET	MOONS	MOON NAMES
Mercury	0	
Venus	0	
Earth	1	Moon
Mars	2	Phobos, Deimos
Jupiter	62	Io, Europa, Ganymede, Callisto, Amalthea, Himalia, Elara, Pasiphae, Sinope, Lysithea, Carme, Ananke, Leda, Metis, Adrastea, Thebe, Callirrhoe, Themisto, Kalyke, Iocaste, Erinome, Harpalyke, Isonoe, Praxidike, Megaclite, Taygete, Chaldene, Autonoe, Thyone, Hermippe, Eurydome, Sponde, Pasithee, Euanthe, Kale, Orthosie, Euporie, Aitne, plus others yet to receive names
Saturn	33	Titan, Rhea, Iapetus, Dione, Tethys, Enceladus, Mimas, Hyperion, Prometheus, Pandora, Phoebe, Janus, Epimetheus, Helene, Telesto, Calypso, Atlas, Pan, Ymir, Paaliaq, Siarnaq, Tarvos, Kiviuq, Ijiraq, Thrym, Skadi, Mundilfari, Erriapo, Albiorix, Suttung, plus others yet to receive names
Uranus	27	Cordelia, Ophelia, Bianca, Cressida, Desdemona, Juliet, Portia, Rosalind, Belinda, Puck, Miranda, Ariel, Umbriel, Titania, Oberon, Caliban, Sycorax, Prospero, Setebos, Stephano, Trinculo, plus others yet to receive names
Neptune	13	Triton, Nereid, Naiad, Thalassa, Despina, Galatea, Larissa, Proteus, plus others yet to receive names
Pluto	1	Charon
<b>TOTAL</b>	<b>139</b>	

Enceladus, Moon of Saturn, by Cassini Orbiter, 2005





# The ~~Nine~~8 Planets

A Multimedia Tour of the Solar System:

one star, eight planets, and more

by Bill Arnett

Pluto demoted,  
to dwarf planet  
2006



Pluto was named  
in 1930 by  
Venetia Burney  
(1918-2009)



<http://www.nineplanets.org/>